Simulating photons, atoms and molecules with quantum phase-space methods

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Simplicity of Photons and Ultracold Gases

- underlying interactions are well understood
- easily characterised by a few parameters
- interactions can be tuned
- → use simple theoretical models to high accuracy
- → develop and test new methods of calculation

Theoretical Methods

- * deterministic methods:
 - → exact diagonalisation X intractable for ≥ 5 particles
 - → factorization X not applicable for strong correlations
 - → perturbation theory X diverges at strong couplings
 - density functional theory X introduces approximations
- probablistic methods:
 - quantum Monte Carlo (QMC)
 - → stochastic wavefunction
 - → phase-space methods

Overview

- introduction to phase-space representations
- density operator description of quantum evolution (3 classes)
 - → static, unitary and open
- Gaussian operator bases (3 types)
 - → coherent, thermal and squeezed
- # applications (3 examples)
 - pulse propagation in optical fibres (photons)
 - → Hubbard model (atoms)
 - → simple atomic-molecular dynamics (molecules)

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Phase-space distributions

- * A classical state can be represented by a joint probability distribution in phase space $P(\mathbf{x}, \mathbf{p})$
- 1932: Wigner constructed an analogous quantity for a quantum state:

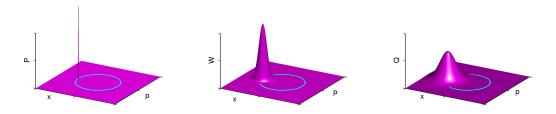
$$W(x,p) = \frac{2}{\pi} \int dy \psi^*(x-y) \psi(x+y) \exp(-2iyp/\hbar)$$

- ★ but it is not always positive → not a true joint probability
- a positive Wigner function would mean a hidden variable interpretation
 of QM is valid

Probability distributions

- * many ways to define phase-space distributions:
 - → eg Wigner, Husimi Q and Glauber-Sudarshan P
 - all defined in terms of coherent states
 - correspond to different choices of orderings
- * to be a probabilistic representation, the phase-space functions must:

	P	W	Q
exist and be nonsingular	X	/	/
always be positive	X	X	/
evolve via drift and diffusion	X	X	X



Reversibility

- * classical random process is irreversible
 - → outward (positive) diffusion
- * quantum mechanics is reversible
 - → phase-space functions generally don't have positive diffusion

A solution!

- * dimension doubling
 - → diffusion into 'imaginary' dimensions ✓
 - → observables evolve reversibly
 - → also fixes up existence and positivity

Phase-space representation

$$\widehat{\rho} = \int P(\overrightarrow{\lambda}) \widehat{\Lambda}(\overrightarrow{\lambda}) d\overrightarrow{\lambda}$$

- $Region{array}{c} P(\overrightarrow{\lambda})$ is a probability distribution
- $\Re \widehat{\Lambda}(\overrightarrow{\lambda})$ is a suitable operator basis
- * is a generalised phase-space coordinate
- * $d\overrightarrow{\lambda}$ is an integration measure
- * equivalent to

$$\widehat{\rho} = E\left[\widehat{\Lambda}(\overrightarrow{\lambda})\right]$$

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Many-body quantum state

Most general description is given by a density operator:



Encapsulates all correlations in a quantum state:

$$\left\langle \widehat{O} \right\rangle = \operatorname{Tr} \left\{ \widehat{O} \, \widehat{\mathsf{p}} \right\}$$

* A probabilistic expansion in the eigenbasis:

$$\widehat{\mathsf{p}} = \sum_{k} P_{k} \ket{\Psi_{k}} ra{\Psi_{k}}$$

Density operators for quantum evolution

1. Unitary dynamics: $\widehat{\rho}(t) = e^{-i\widehat{H}t/\hbar}\widehat{\rho}(0)e^{i\widehat{H}t/\hbar}$

2. Equilibrium state: $\widehat{\rho}_{\rm un}(T) = e^{-(\widehat{H} - \mu \widehat{N})/k_B T}$

3. Open dynamics: $\widehat{\rho}_{Sys} = Tr_{Res} \{\widehat{\rho}\}$

* each type is equivalent to a Liouville equation for $\hat{\rho}$:

$$\frac{d}{d\tau}\widehat{\rho} = \widehat{L}[\widehat{\rho}]$$

Phase-space Recipe

- 1. Formulate: $\partial \widehat{\rho}/\partial \tau = \widehat{L}[\widehat{\rho}]$
- 2. Expand: $\int \partial P/\partial \tau \widehat{\Lambda} d\overrightarrow{\lambda} = \int P\widehat{L} \left[\widehat{\Lambda} \right] d\overrightarrow{\lambda}$
- 3. Transform: $\widehat{L}\left[\widehat{\Lambda}\right] = L\widehat{\Lambda}$
- 4. Integrate by parts: $\int P \mathcal{L} \widehat{\Lambda} d\overrightarrow{\lambda} \Longrightarrow \int \widehat{\Lambda} \mathcal{L}' P d\overrightarrow{\lambda}$
- 5. **Obtain** Fokker-Planck equation: $\partial P/\partial \tau = \mathcal{L}'P$
- 6. **Sample** with stochastic equations for $\overrightarrow{\lambda}$

Stochastic Gauges

- Mapping from Hilbert space to phase space not unique
 - → many "gauge" choices
- * Can alter noise terms B_{ij} , introduce arbitrary drift functions $g_j(\overrightarrow{\lambda})$

Weight
$$d\Omega/d\tau = \Omega \left[U + g_j \zeta_j\right]$$

Trajectory
$$d\lambda_i/\partial \tau = A_i + B_{ij}[\zeta_j - g_j]$$

Interacting many-body physics

$$\widehat{\rho} \Longrightarrow \overrightarrow{\lambda}$$

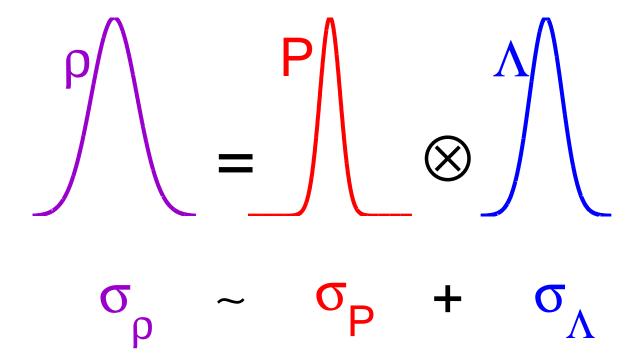
- many-body problems map to nonlinear stochastic equations
- calculations can be from first-principles
- ✓ precision limited only by sampling error
- ✓ choose basis to suit the problem

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Operator Bases

* need basis simple enough to fit into a computer, complex enough to contain the relevant physics:



General Gaussian operators

a generalisation of the density operators that describe Gaussian states

- ***** Gaussian states can be:
 - → coherent (for bosons), squeezed, or thermal







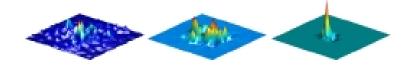
- → or any combination of these
- * characterised by first-order moments: \overline{x} , \overline{p} , $\overline{x^2}$, $\overline{p^2}$, \overline{xp}
 - → all higher-order moments factorise

Gaussian Basis I: Coherent-state projectors

$$\widehat{\Lambda} = \frac{|\alpha\rangle\langle\beta|}{\langle\beta||\alpha\rangle}$$

- lpha defines the +P distribution, with a doubled phase space $\overrightarrow{\lambda}=(\Omega,\alpha,\beta)$
- * moments: $\langle O(\widehat{\mathbf{a}}^{\dagger}, \widehat{\mathbf{a}}) \rangle = E[O(\beta^*, \alpha)]$
- * successful for many applications in quantum optics
- * successful simulations of short-time quantum dynamics of BEC

Evaporative Cooling of a BEC



- # first-principles 3D calculation
 - \rightarrow start with Bose gas above T_c ; finish with narrow BEC peak
 - → 20000 atoms, 32000 modes
 - → Hilbert space is astronomically large
- X Problems!
 - method pushed to the limit
 - * breaks down for longer times, stronger interactions

Gaussian Basis II: Thermal operators

$$\widehat{\mathbf{\Lambda}} = |\mathbf{I} \pm \mathbf{n}|^{\mp 1} : \exp \left[\widehat{\mathbf{a}} \left(\mathbf{I} \mp \mathbf{I} - \left[\mathbf{I} \pm \mathbf{n} \right]^{-1} \right) \widehat{\mathbf{a}}^{\dagger} \right] :$$

- lpha now have a phase-space of variances: $\overrightarrow{\lambda} = (\Omega, \mathbf{n})$
- * defined for bosons (upper sign) and fermions (lower sign)
- lpha moments: $\left\langle \widehat{a}_{i}^{\dagger}\widehat{a}_{j}\right\rangle =E\left[n_{ij}\right]$, $\left\langle \widehat{a}_{i}^{\dagger}\widehat{a}_{j}^{\dagger}\widehat{a}_{j}\widehat{a}_{i}\right\rangle =E\left[n_{ii}n_{jj}-n_{ij}n_{ji}\right]$
- suitable for cold atoms

Gaussian Basis III: General form (including squeezing)

$$\widehat{\Lambda}(\overrightarrow{\lambda}) = \Omega \sqrt{\left|\underline{\underline{\sigma}}\right|^{\mp 1}} : \exp\left[\delta \widehat{\underline{a}}^{\dagger} \left(\underline{\underline{I}} \mp \underline{\underline{I}} - \underline{\underline{\sigma}}^{-1}\right) \delta \widehat{\underline{a}}/2\right] :$$

relative displacement: $\delta \underline{\hat{a}} = \underline{\hat{a}} - \underline{\alpha}$

annihilation and creation operators: $\underline{\widehat{a}} = \left(\widehat{a}_1,...,\widehat{a}_M,\widehat{a}_1^\dagger,...,\widehat{a}_M^\dagger\right)$

coherent offset: $\underline{\alpha} = (\alpha_1, ..., \alpha_M, \alpha_1^+, ..., \alpha_M^+)$, $(\underline{\alpha} = 0$ for fermions)

upper signs: bosons; lower signs: fermions

Extended phase space

$$\overrightarrow{\lambda} = (\Omega, \alpha, \alpha^+, \mathbf{n}, \mathbf{m}, \mathbf{m}^+)$$

- \longrightarrow Hilbert-space dimension: 2^M for fermions, N^M for bosons
- \implies phase-space dimension: $2(1-M+2M^2)$ for fermions, $2(1+3M+2M^2)$ for bosons

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Application I: photons in a fibre

$$\widehat{H} = \widehat{H}_F + \widehat{H}_L + \widehat{H}_G + \widehat{H}_R$$

- * \widehat{H}_F : fibre-optic Hamiltonian, including $\chi^{(3)}$ nonlinearity
- lpha \widehat{H}_L , \widehat{H}_G : coupling to absorbing reserviors and fibre amplifier reserviors
- $\# \widehat{H}_R$: nonlinear coupling to non-Markovian phonon reserviors
- * have 10^2 modes and 10^9 particles

Phase-Space Equations

use coherent-state basis → get a stochastic, Raman modified NLS equations:

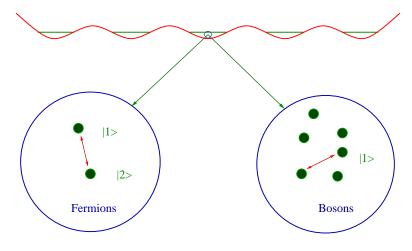
$$\frac{\partial}{\partial x} \phi(t, x) = -\int_{-\infty}^{\infty} dt' g(t - t') \phi(t', x) + \Gamma(t, x) \pm \frac{i}{2} \frac{\partial^2}{\partial t^2} \phi$$

$$+ \left[i \int_{-\infty}^{\infty} dt' h(t - t') \phi^+(t', x) \phi(t', x) + \Gamma^R(t, x) \right] \phi(t, x)$$

$$\frac{\partial}{\partial x} \phi^+(t, x) = -\int_{-\infty}^{\infty} dt' g^*(t - t') \phi^+(t', x) + \Gamma^+(t, x) \mp \frac{i}{2} \frac{\partial^2}{\partial t^2} \phi$$

$$+ \left[i \int_{-\infty}^{\infty} dt' h^*(t - t') \phi(t', x) \phi^+(t', x) + \Gamma^{+R}(t, x) \right] \phi^+(t, x)$$

Application II: atoms in a lattice



$$\widehat{H} = -t \sum_{\langle i,j \rangle,\sigma} \widehat{a}_{i,\sigma}^{\dagger} \widehat{a}_{j,\sigma} + U \sum_{j} \widehat{a}_{j,\uparrow}^{\dagger} \widehat{a}_{j,\downarrow}^{\dagger} \widehat{a}_{j,\downarrow} \widehat{a}_{j,\downarrow} \widehat{a}_{j,\uparrow}$$

- * simplest model of an interacting Fermi gas on a lattice
 - → weak-coupling limit → BCS transitions
 - \rightarrow solid-state models; relevance to High- T_c superconductors

Solving the Hubbard Model

- * only the 1D model is exactly solvable (Lieb & Wu, 1968)
- * even then, not all correlations can be calculated
- higher dimensions can use Quantum Monte Carlo methods.
- ★ except for a few special symmetrical cases, QMC suffers from sign problems with the Hubbard model
 - * e.g. sign problems for repulsive interaction away from half filling
- x sign problem increases with dimension, lattice size, interaction strength

Fermionic sign problem

- ** Quantum Monte Carlo (QMC) samples many-body wavefunction $\phi(r)$ (wavefunction treated as a probability)
- * but Fermion states are antisymmetric
 - → wavefunction nonpositive
- must introduce (possibly negative) weighting factors
 - → bad sampling errors (unless approximations used)

$$\langle A \rangle \sim \frac{\langle sA \rangle}{\langle s \rangle}$$

Applying the Gaussian representation

Use thermal basis, and apply mappings

$$\begin{array}{lcl} \widehat{\mathbf{n}}_{\sigma}\widehat{\boldsymbol{\rho}} & \rightarrow & \left\{ 2\mathbf{n}_{\sigma} - (\mathbf{I} - \mathbf{n}_{\sigma}) \frac{\partial}{\partial \mathbf{n}_{\sigma}} \mathbf{n}_{\sigma} \right\} P(\Omega, \mathbf{n}_{\parallel}, \mathbf{n}_{\parallel}) \\ \\ \widehat{\boldsymbol{\rho}}\widehat{\mathbf{n}}_{\sigma} & \rightarrow & \left\{ 2\mathbf{n}_{\sigma} - \mathbf{I} - \mathbf{n}_{\sigma} \frac{\partial}{\partial \mathbf{n}_{\sigma}} (\mathbf{I} - \mathbf{n}_{\sigma}) \right\} P(\Omega, \mathbf{n}_{\parallel}, \mathbf{n}_{\parallel}) \\ \\ \widehat{\boldsymbol{\rho}} & \rightarrow & - \frac{\partial}{\partial \Omega} \Omega(\Omega, \mathbf{n}_{\parallel}, \mathbf{n}_{\parallel}) \end{array}$$

- \implies Fokker-Planck equation for P, with drift and diffusion
- \Longrightarrow sample with stochastic equations for Ω and \mathbf{n}_{σ}

Positive-Definite Diffusion

Modify interaction term with a 'Fermi gauge':

$$U\sum_{j}:\widehat{n}_{jj,\downarrow}\widehat{n}_{jj,\uparrow}: = -\frac{1}{2}|U|\sum_{j}:\left(\widehat{n}_{jj,\downarrow}-\frac{U}{|U|}\widehat{n}_{jj,\uparrow}\right)^{2}:$$

- diffusion matrix has a real 'square root' matrix
 - realise the diffusion with a real noise process
 - problem maps to a real (and much more stable) subspace

Stratonovich Equations

* Stratonovich stochastic equations, in matrix form:

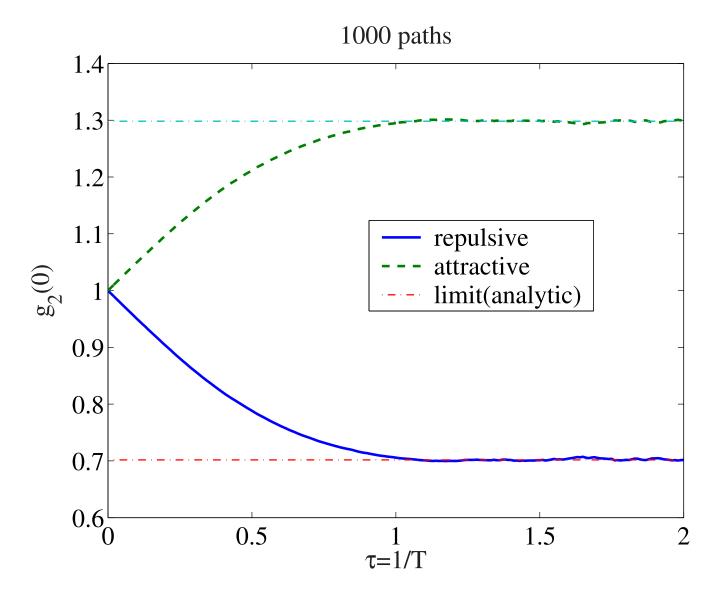
$$\frac{d\Omega}{d\tau} = -\Omega \left\{ -t \sum_{\langle i,j\rangle,\sigma} n_{ij,\sigma} + U \sum_{j} n_{jj,\downarrow} n_{jj,\uparrow} - \mu \sum_{j,\sigma} n_{jj,\sigma} \right\}
\frac{d\mathbf{n}_{\sigma}}{d\tau} = -\frac{1}{2} \left\{ (\mathbf{I} - \mathbf{n}_{\sigma}) \Delta_{\sigma}^{(1)} \mathbf{n}_{\sigma} + \mathbf{n}_{\sigma} \Delta_{\sigma}^{(2)} (\mathbf{I} - \mathbf{n}_{\sigma}) \right\},$$

where the stochastic propagator matrix is

$$\Delta_{i,j,\sigma}^{(r)} = \left[-t_{ij} + \delta_{i,j} \left\{ U n_{jj,\sigma'} - |U| n_{jj,\sigma} + \frac{1}{2} |U| - \mu \right\} \right] \pm \delta_{ij} \sqrt{2|U|} \xi_j^{(r)}$$

 $*\xi_{j}^{(r)}$ are delta-correlated white noises

1D Lattice-100 sites



Branching

* averages are weighted,eg

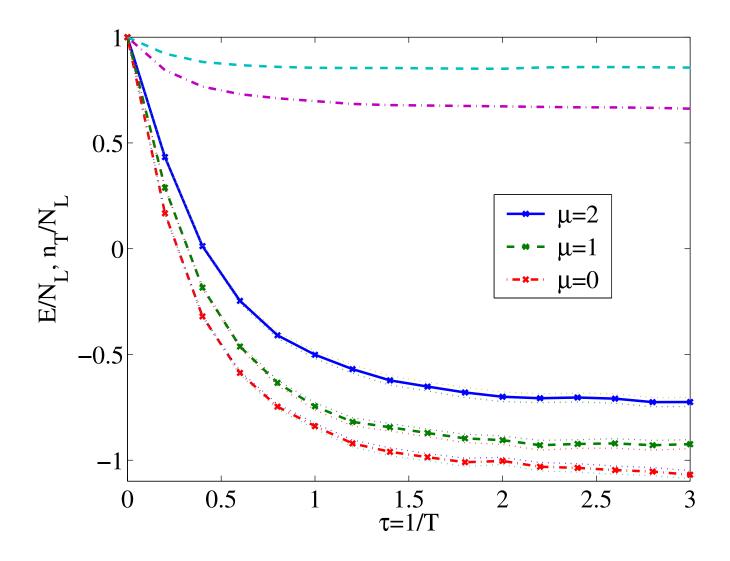
$$raket{\widehat{\mathbf{n}}(\mathbf{ au})} = rac{\sum_{j=1}^{N_p} \mathbf{\Omega}^{(j)}(\mathbf{ au}) \mathbf{n}^{(j)}(\mathbf{ au})}{\sum_{j=1}^{N_p} \mathbf{\Omega}^{(j)}(\mathbf{ au})}$$

- ✗ but weights spread exponentially → many irrelevant paths
- delete low-weight paths and clone high-weight paths:

$$m^{(jp)} = \operatorname{Integer}\left[\xi + \Omega^{(jp)}/\overline{\Omega}\right]$$

- $*\xi \in [0,1]$ is a random variable, $\overline{\Omega}$ is an average weight
- * after branching, weights of surviving paths are equalised

16x16 2D Lattice

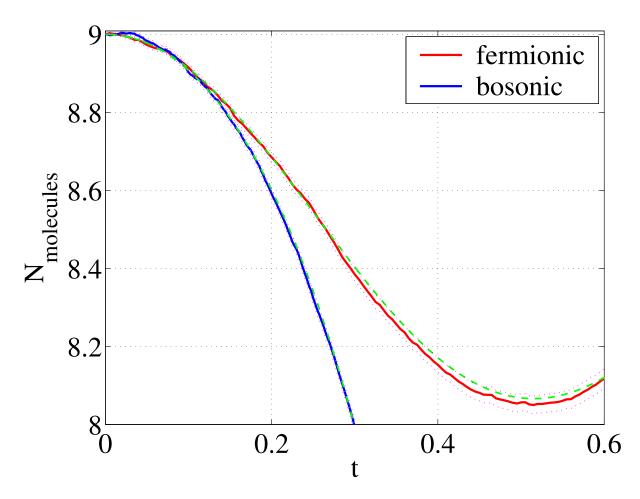


* No sign problem!

Application III: Molecules in a well

lpha Hamiltonian: $\widehat{H}=\widehat{a}\widehat{b}_1^{\dagger}\widehat{b}_2^{\dagger}+\widehat{a}^{\dagger}\widehat{b}_1\widehat{b}_2$

Result: Pauli blocking



Summary

- * Generalised phase-space representations provide a means of simulating many-body quantum physics from first principles, with *precision limited only by sampling error*.
- ** Coherent-state-based methods have been successful in simulating quantum dynamics of photons and weakly interacting ultracold gases.
- * Gaussian-based methods extend the applicability to highly correlated sysems of bosons and *fermions*.
- * Simulated the Hubbard model (fermions in a lattice) without sign errors.